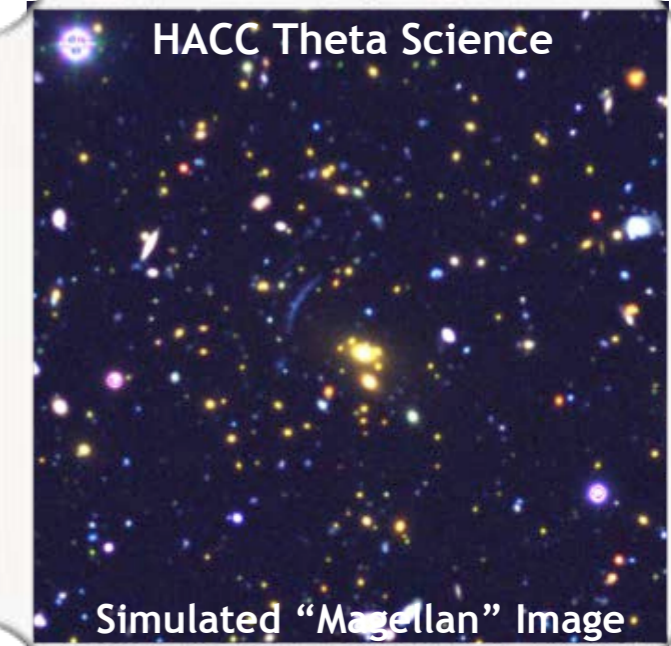
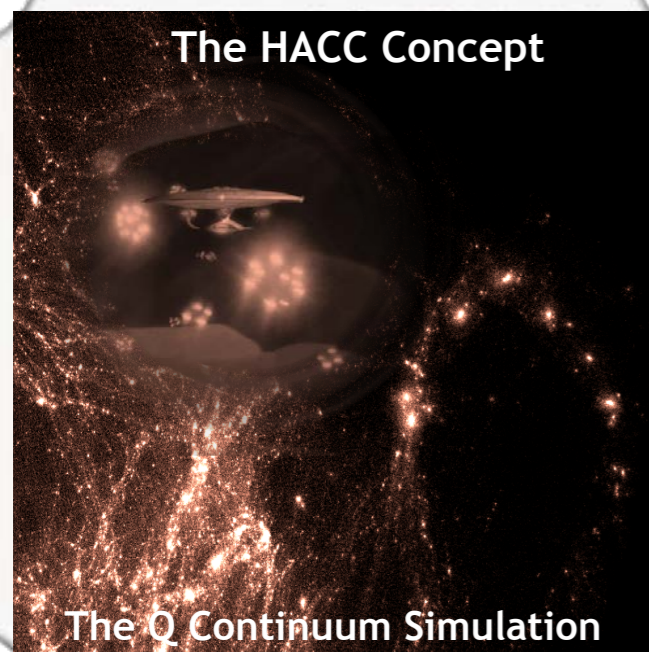


Next-Generation Cosmology Simulations with HACC: Challenges from Baryons

Katrin Heitmann

Lunch Talk, Computational Performance Workshop



* HACC = Hardware/Hybrid Accelerated Cosmology Code

The Cosmologist's View on Baryons



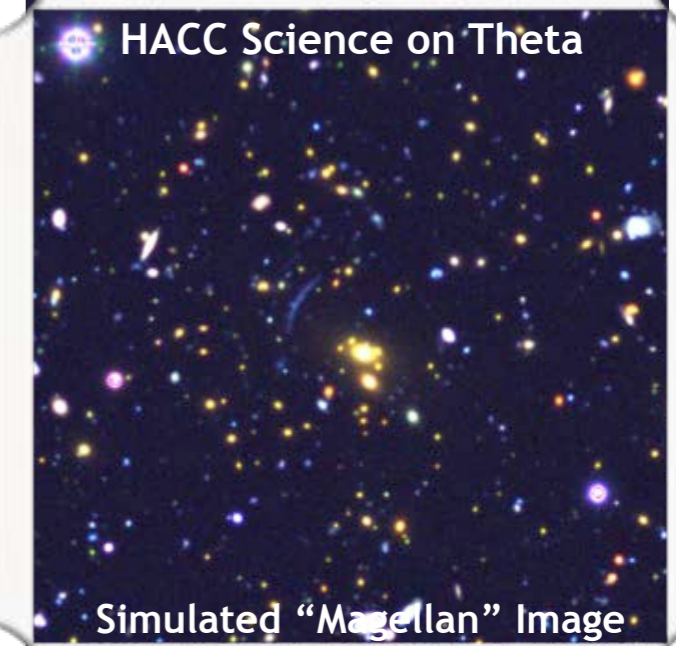
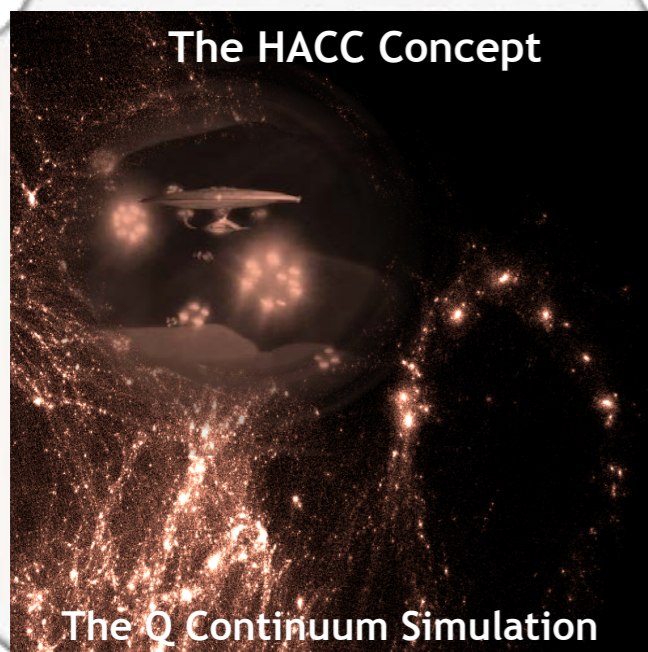
The Cosmologist's View on Baryons



Next-Generation Cosmology Simulations with HACC*: Challenges from Baryons

Katrin Heitmann

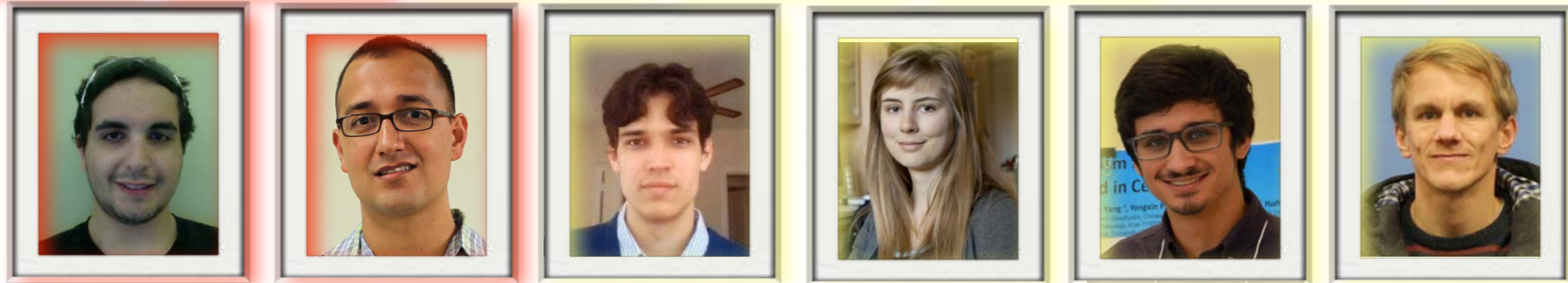
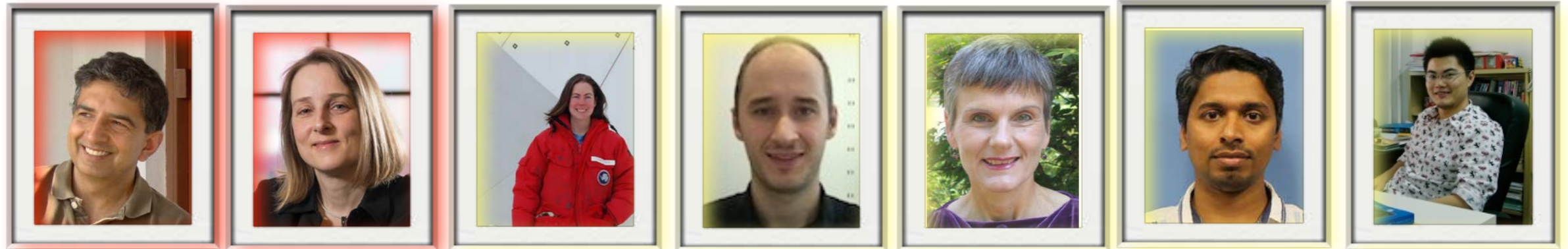
Lunch Talk, Computational Performance Workshop



* HACC = Hardware/Hybrid Accelerated Cosmology Code

Our Group at Argonne and Close Collaborators

HEP Division



ALCF Division



External



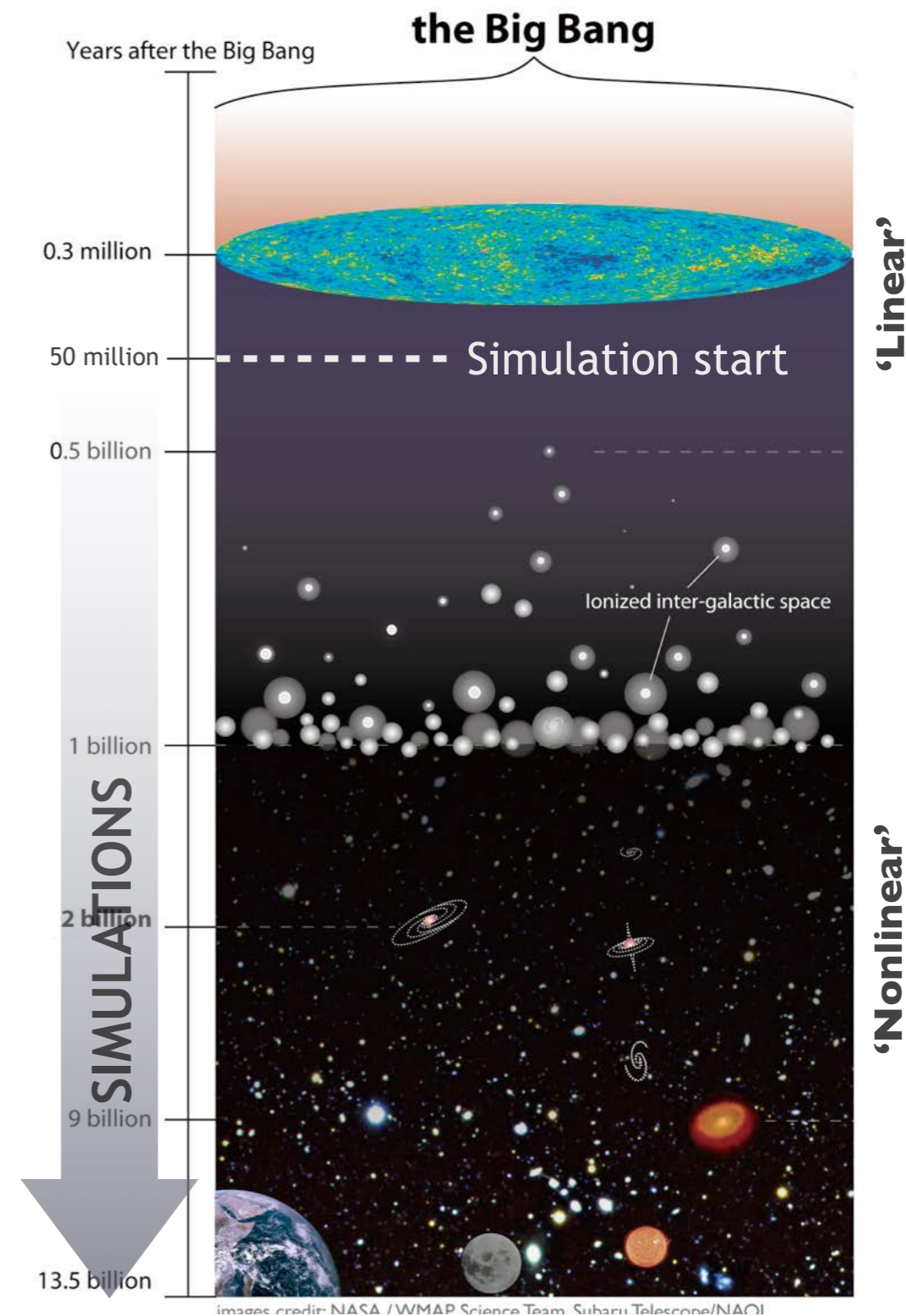
● Core HACC Team

● HACC Science Collaborators



The Evolution of the Universe: Structure Formation

- Solid understanding of structure formation; success underpins most cosmic discovery
 - ▶ Initial conditions determined by primordial fluctuations, measured from the cosmic microwave background
 - ▶ Initial perturbations amplified by gravitational instability in a dark matter-dominated Universe
 - ▶ Relevant theory is gravity, field theory, and atomic physics ('first principles')
- Early Universe: **Linear** perturbation theory very successful (CMB)
- Latter half of the history of the Universe: **Nonlinear** domain of structure formation, **impossible** to treat without large-scale computing



Computing the Universe: Gravity-only

- Gravity dominates on large scales, use Monte Carlo sampling of density with tracer particles
- Particles are tracers of the dark matter in the Universe, mass typically at least $\sim 10^9 M_\odot$

$$m_p \sim V/n_p$$

- Simulate galaxy size objects ($v^2/c^2 \ll 1$), Newtonian description accurate
- At smaller scales, add gas physics, feedback etc., sub-grid modeling inevitable

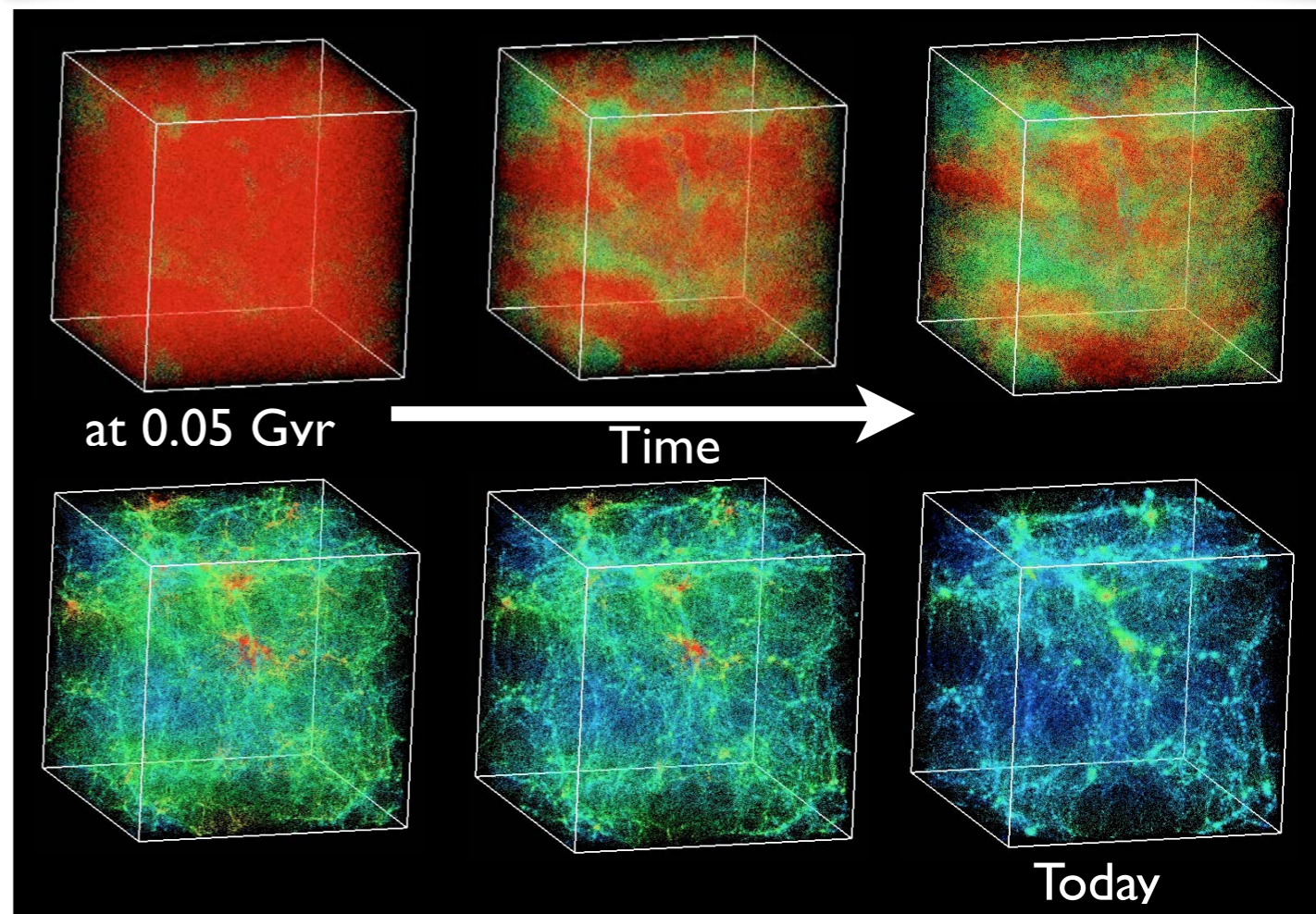
“The Universe is far too complicated a structure to be studied deductively, starting from initial conditions and solving the equations of motion.”

Robert Dicke (Jayne Lectures, 1969)

$$\ddot{\mathbf{x}} + 2\frac{\dot{a}}{a}\dot{\mathbf{x}} = -\frac{\nabla\Phi}{a^2} \quad \text{Equation of motion for tracer particles in expanding Universe}$$

$$\frac{\dot{a}}{a} = H = \frac{H_0}{a^{3/2}} \sqrt{\Omega_{tot} + a^3\Omega_\Lambda} \quad \text{CDM + baryons + DE}$$

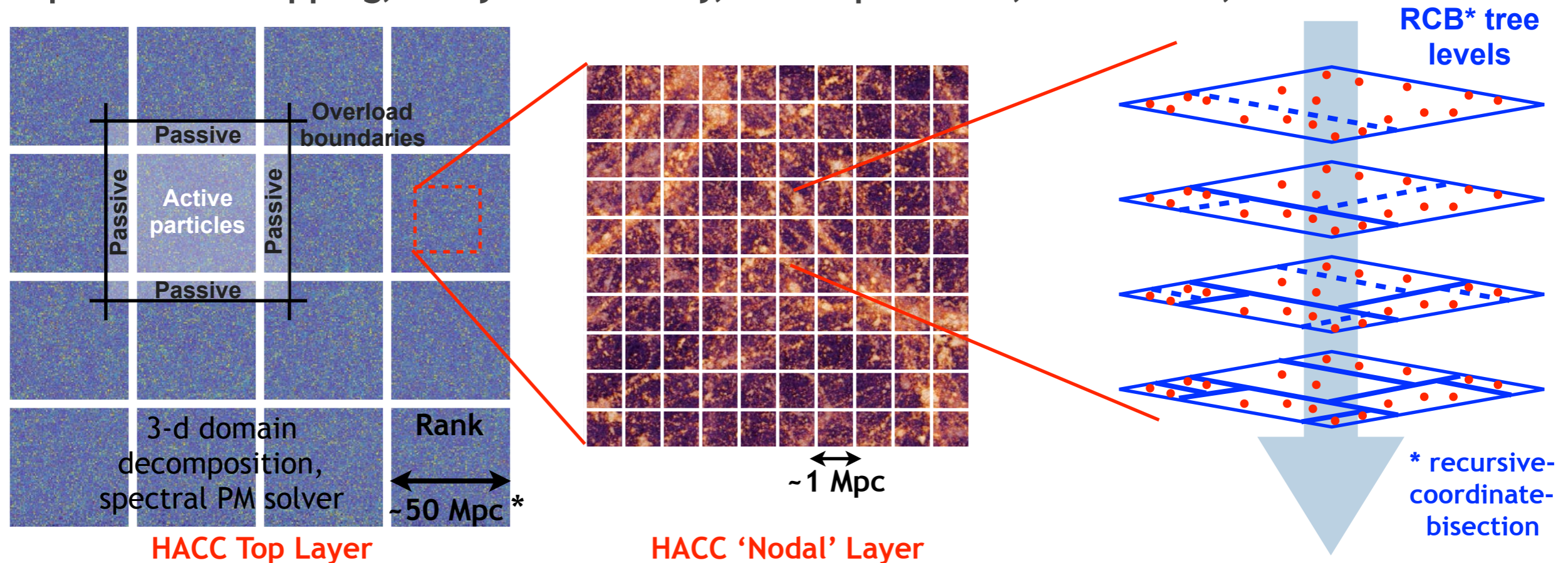
$$\nabla^2\Phi(\mathbf{x}) = 4\pi G a^2 [\rho(\mathbf{x}, t) - \rho_b(t)] \quad \text{Poisson equation}$$



HACC in a Nutshell

S. Habib et al. 2016, New Astronomy

- Long-range/short range force splitting:
 - ▶ **Long-range:** Particle-Mesh solver, C/C++/MPI, **unchanged for different architectures**, FFT performance dictates scaling (custom pencil decomposed FFT)
 - ▶ **Short-range:** **Depending on node architecture** switch between tree and particle-particle algorithm; tree needs “thinking” (building, walking) but computationally less demanding (BG/Q, X86, KNL), PP easier but computationally more expensive (GPU)
- Overload concept to allow for **easy swap of short-range solver** and minimization of communication (reassignment of passive/active in regular intervals)
- Adaptive time stepping, analysis on the fly, mixed precision, custom I/O, ...

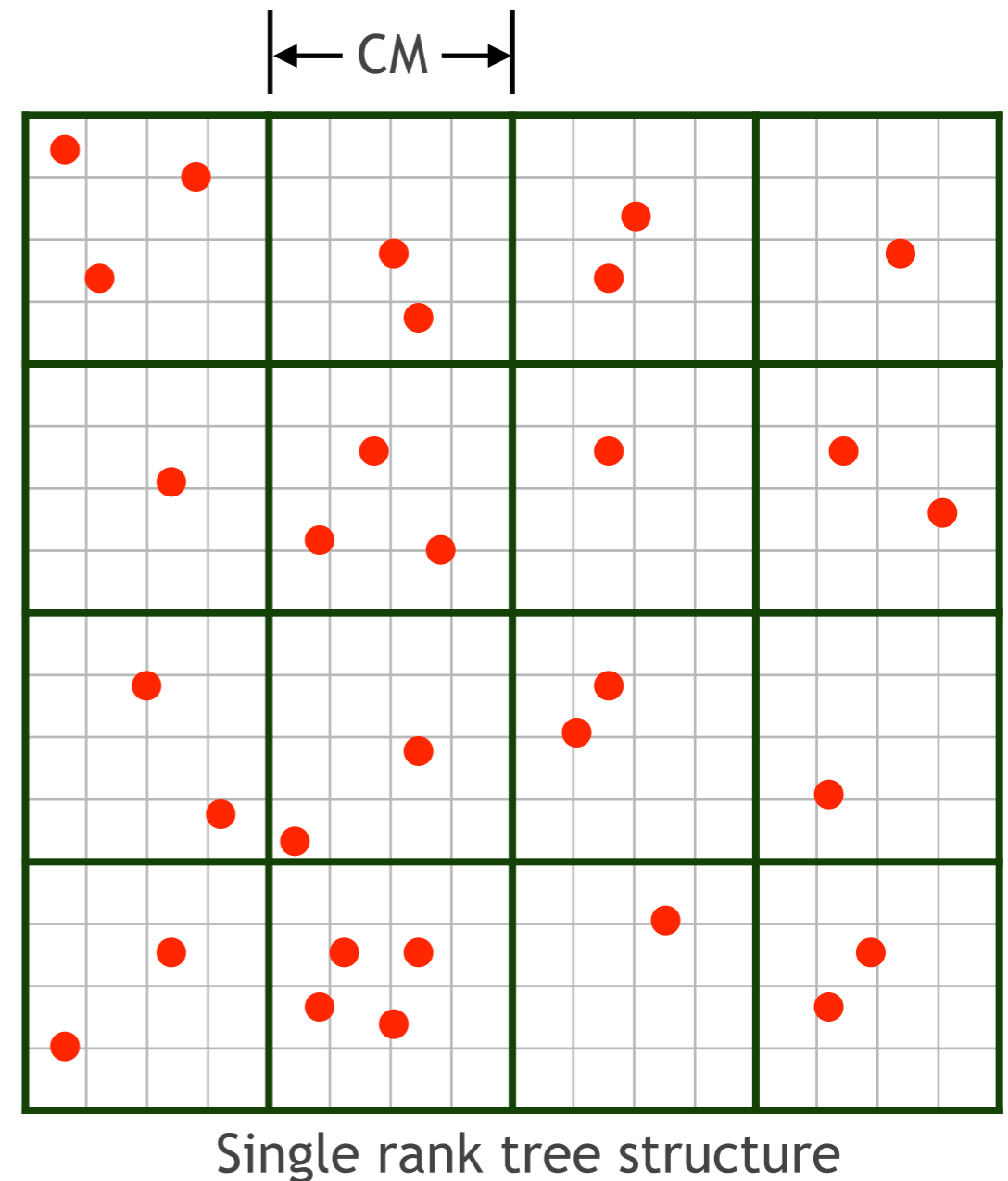


* Mpc $\sim 3M$ light years, distance between 2 galaxies



Tree Portability

- Each rank divides its volume into chaining mesh cells of width **CM**
- An RCB tree is built in each CM cell. Tree bisection stops once the number of particles on a leaf drops below **PPN**
- Customize PP force kernel for specific architecture
- Data structure arranged for **easy vectorization** of the PP kernel and **tree walk parallelized using dynamic OpenMP threads**
- **Flexibility in CM and PPN** allows for **easy portability** to new machines
 - ▶ For Mira (BG/Q): PPN ~ 128 optimal
 - ▶ For Edison (X86): PPN ~ 20 optimal
 - ▶ For Theta (KNL): PPN again higher



HACC Timing: One BG/Q Node versus 1/4 KNL Node

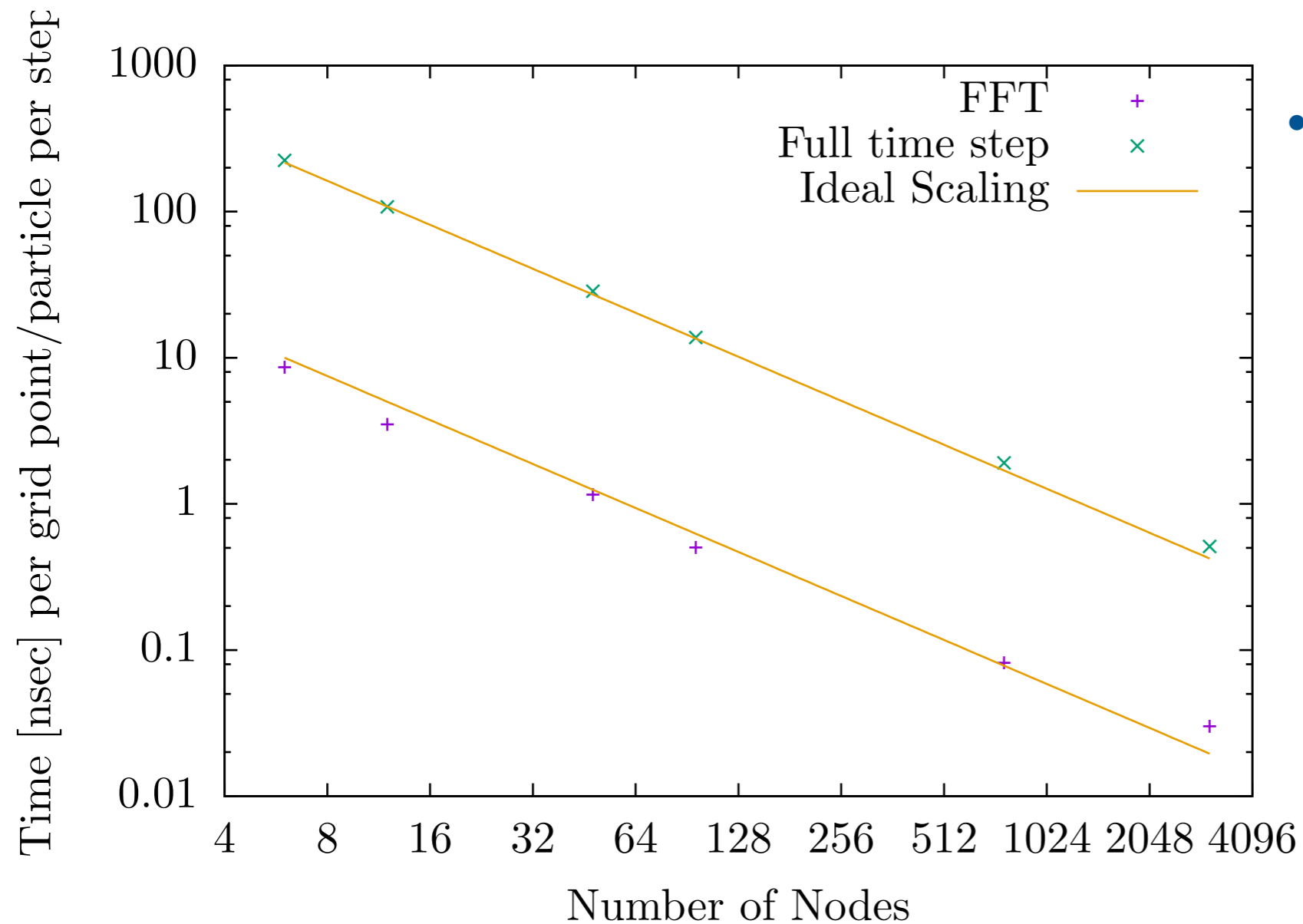
Cores	RPN	OMP	TH	BG/Q [Time/s]	KNL [Time/sec]	Ratio
16	4	4	16	4297	616.33	6.98
16	4	8	32	2677	543.73	4.92
16	4	16	64	2504	530.23	4.72
16	8	2	16	4362	544.75	8.00
16	8	4	32	2571	459.53	5.59
16	8	8	64	2278	437.21	5.18
16	16	4	64	2581	468.50	5.50

RPN: MPI ranks for each run; **OMP:** OpenMP threads per MPI rank, **TH:** total number of threads = RPN*OMP

- Problem set up: 320^3 particles, 320^3 grid, 3 full time steps, 5 sub-cycles
- Results for flat mode and cache mode are essentially identical



HACC Weak Scaling on Theta

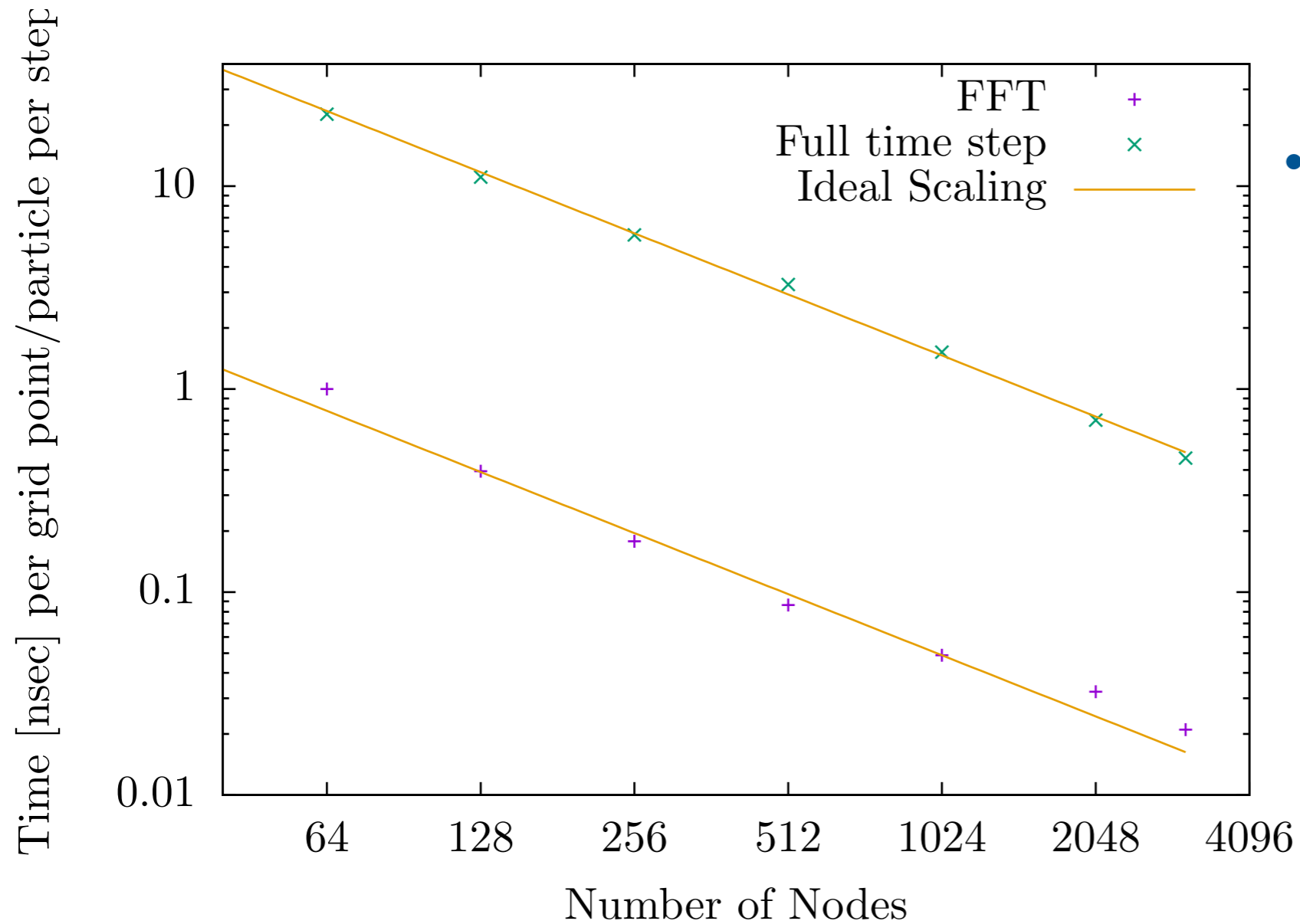


- **Problem set up:**

- ▶ 6 nodes: 1152^3 particles in $(312.5 \text{ Mpc/h})^3$ volume
- ▶ 3072 nodes: 9216^3 particles in $(2500 \text{ Mpc/h})^3$ volume
- ▶ 8 ranks per nodes, 16 threads
- ▶ Cache-quad mode



HACC Strong Scaling on Theta



- **Problem set up:**

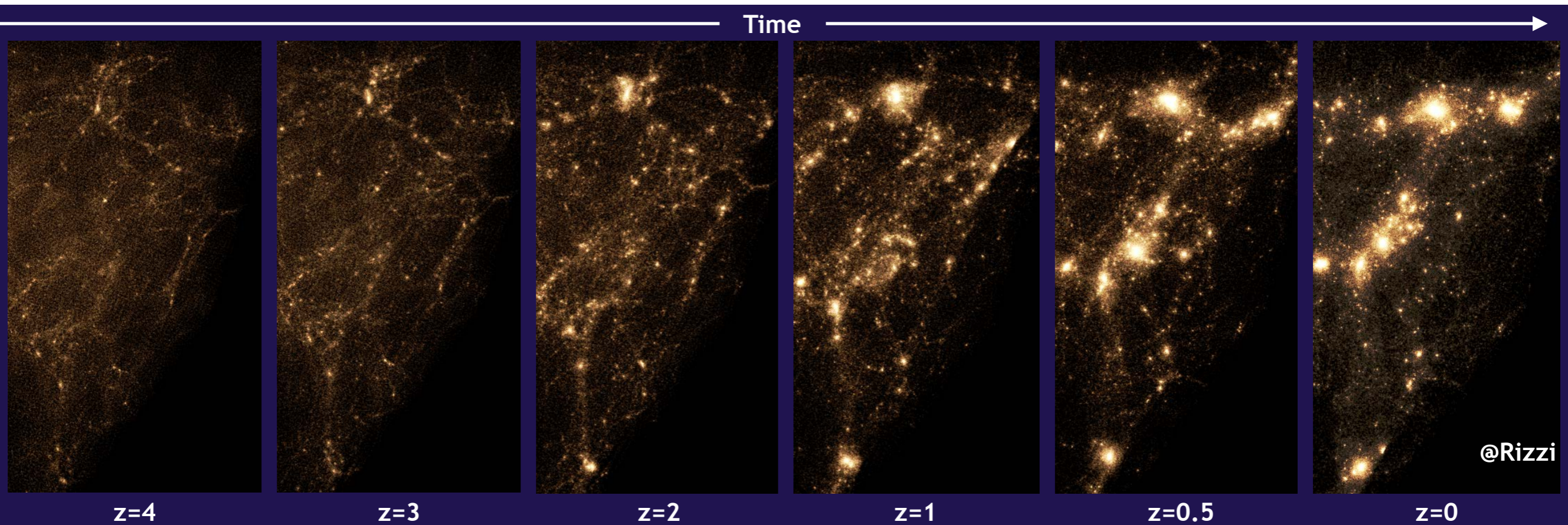
- ▶ 2048³ particles on 64 to 2048 nodes (and 2304³ particles on 3072 nodes)
- ▶ 8 ranks per nodes, 16 threads
- ▶ Cache-quad mode



The Borg Cube Simulation

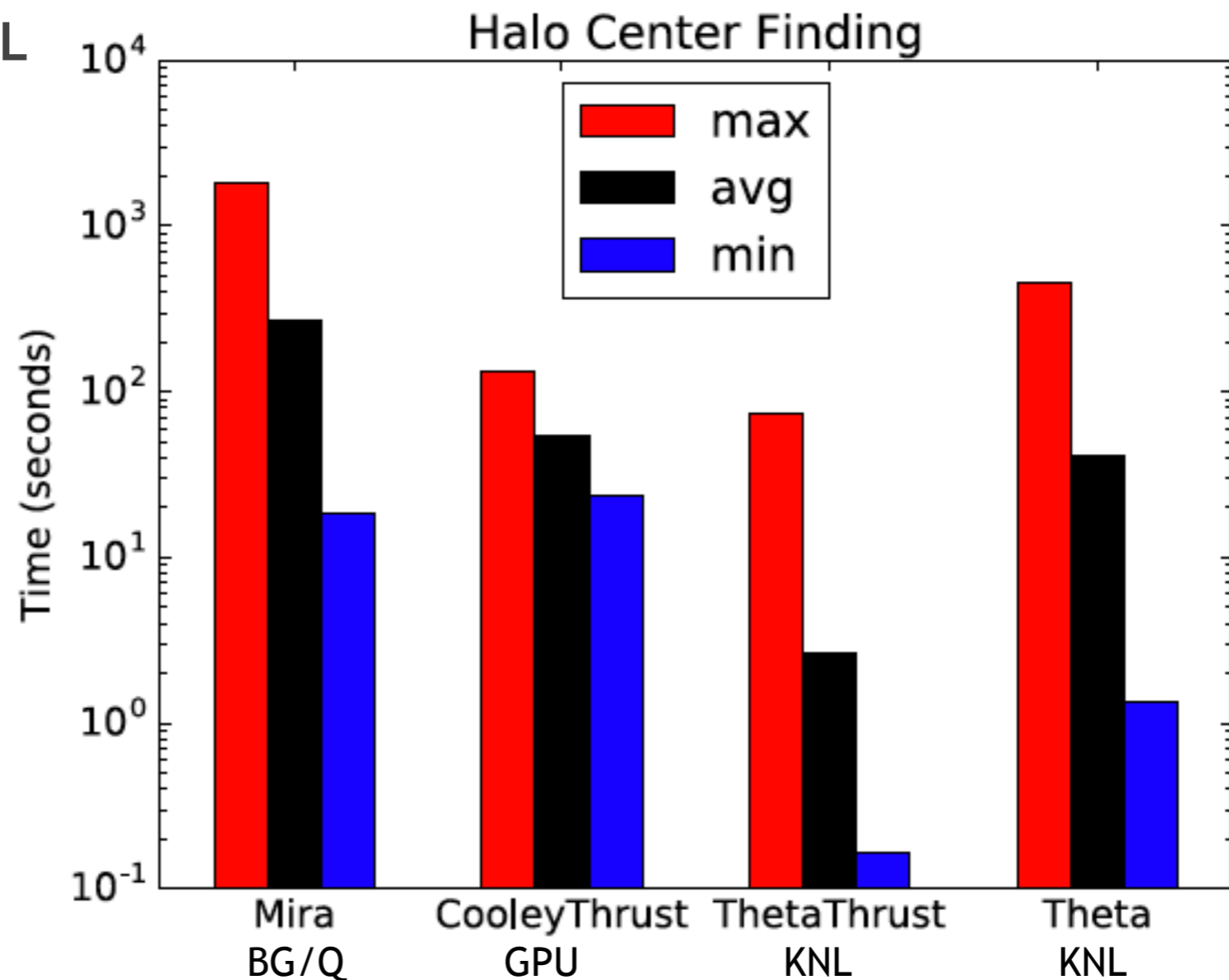
- **Science target on Theta**

- ▶ $(2 \times 2304)^3$ particles in $(800 \text{ Mpc}/h)^3$ volume
- ▶ Study of maps relevant to cross-correlation between cosmic microwave background and large-scale structure
- ▶ Gravity-only simulation + modeling of baryons in post-processing vs. hydro simulation
- ▶ Gravity-only simulation finished

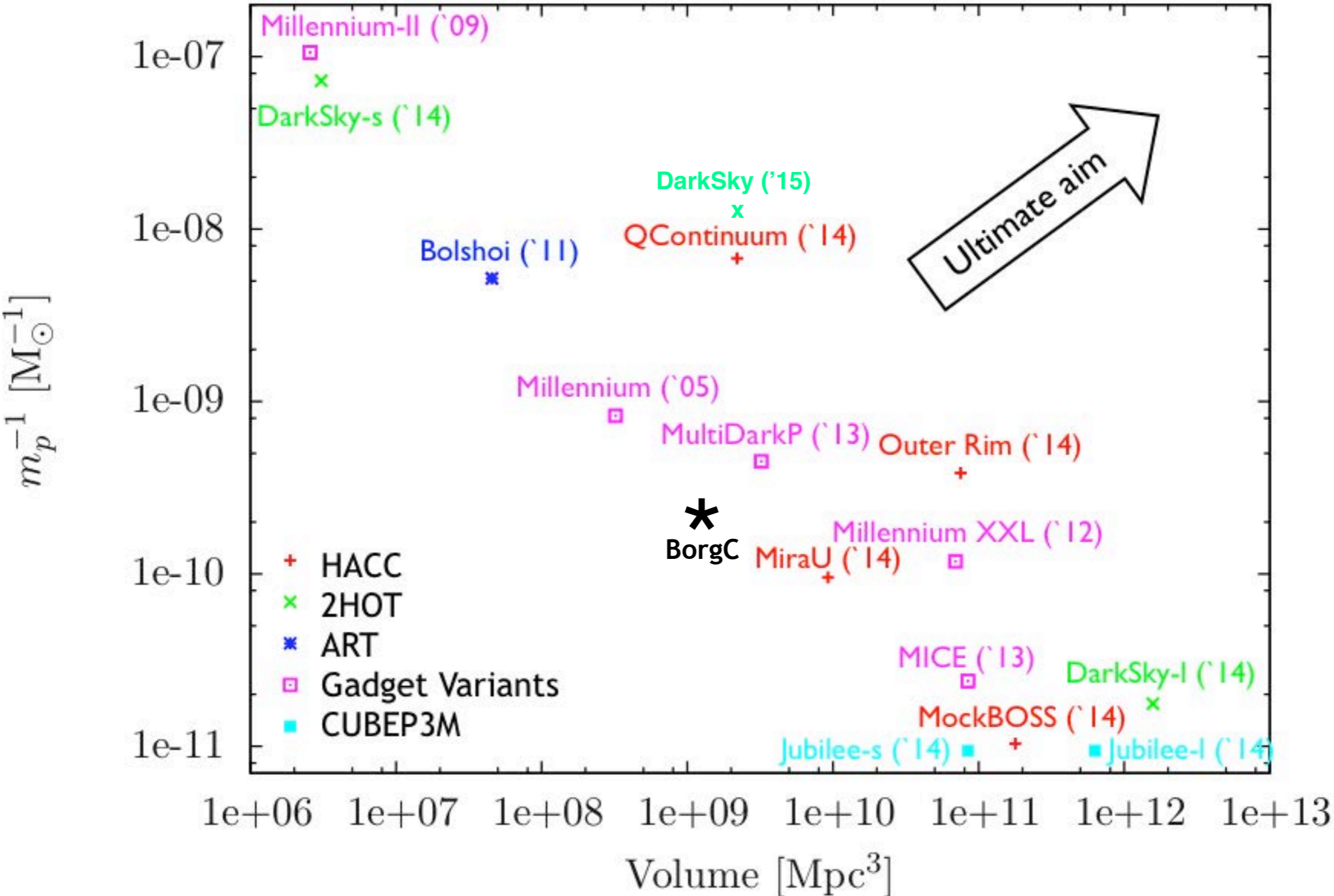


Portable Analysis Tools with Thrust

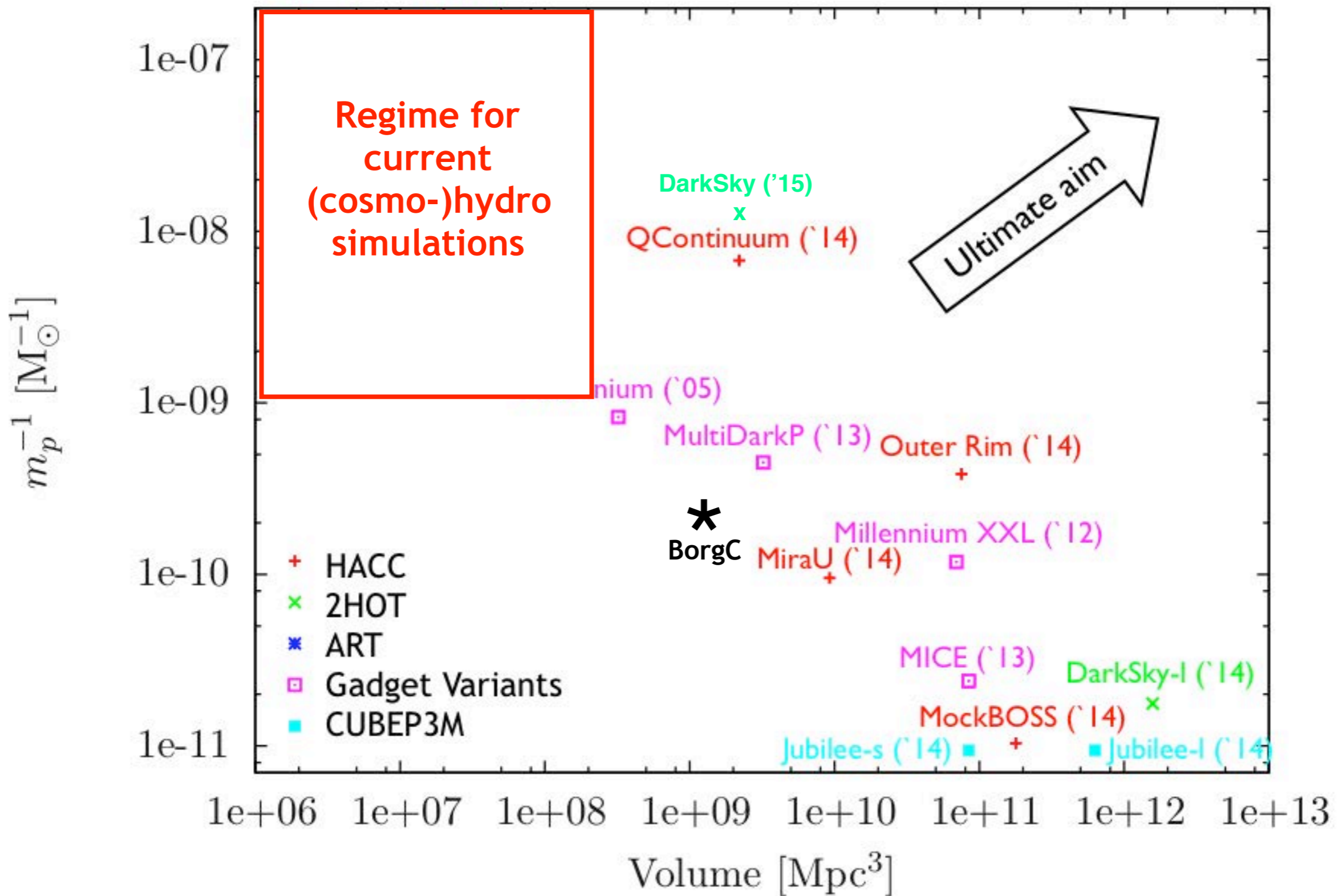
- Halo finder identifies groups of particles within a certain distance
- Center finder is expensive piece, calculate local potential minimum
- Thrust: NVIDIA library of parallel algorithms and data structures to perform sorts, scans, transforms, etc.
- Works on GPU and KNL
- Test on 64 KNL nodes



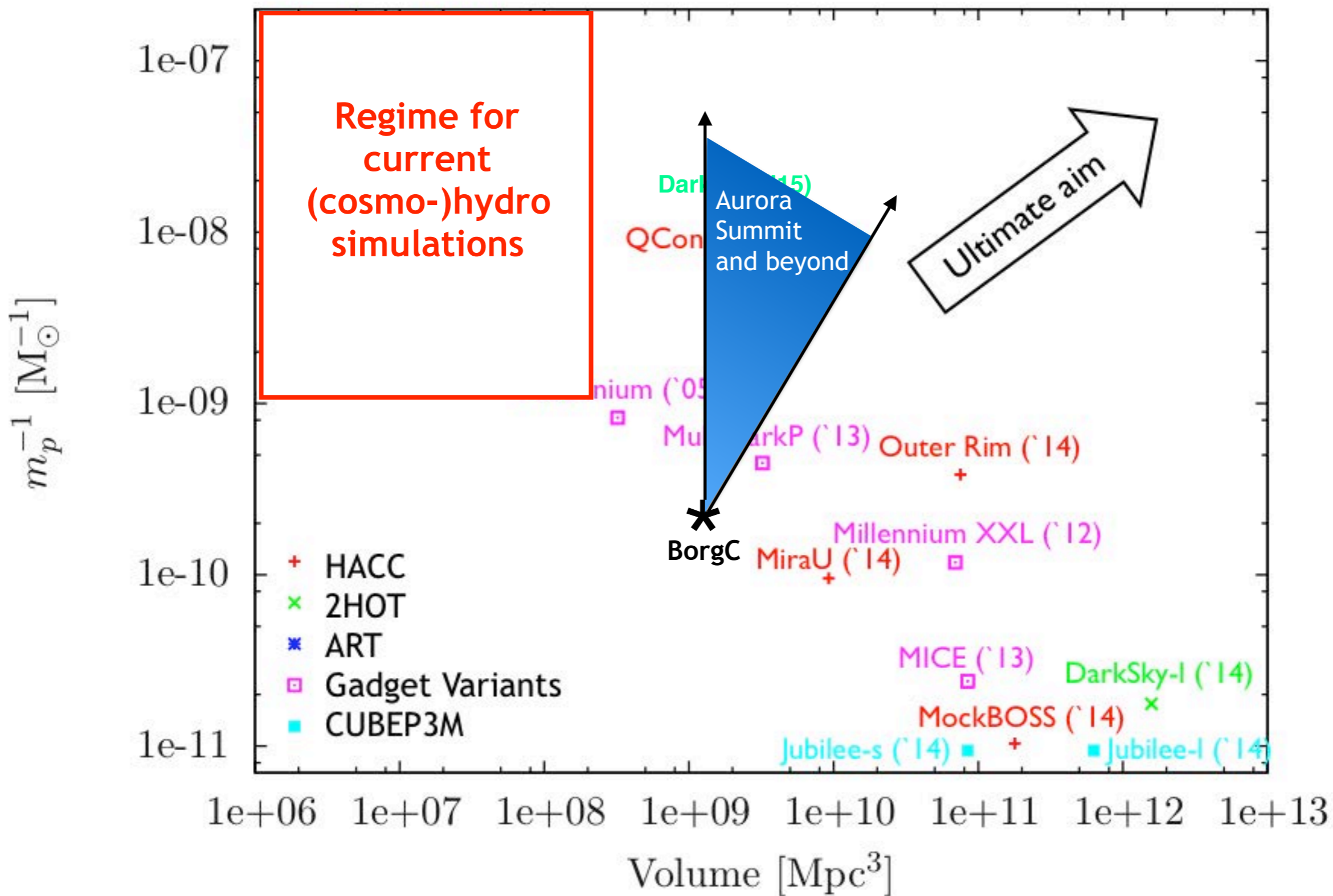
Current State-of-the-Art Gravity-Only Simulations



Current State-of-the-Art Simulations

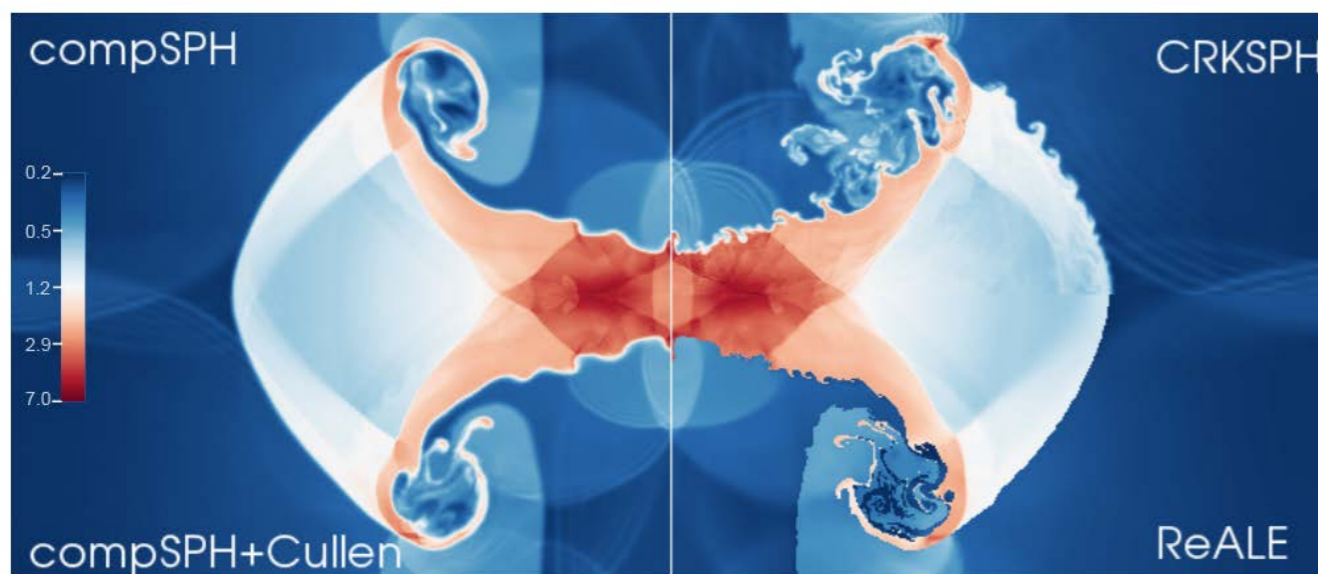


Current State-of-the-Art Simulations

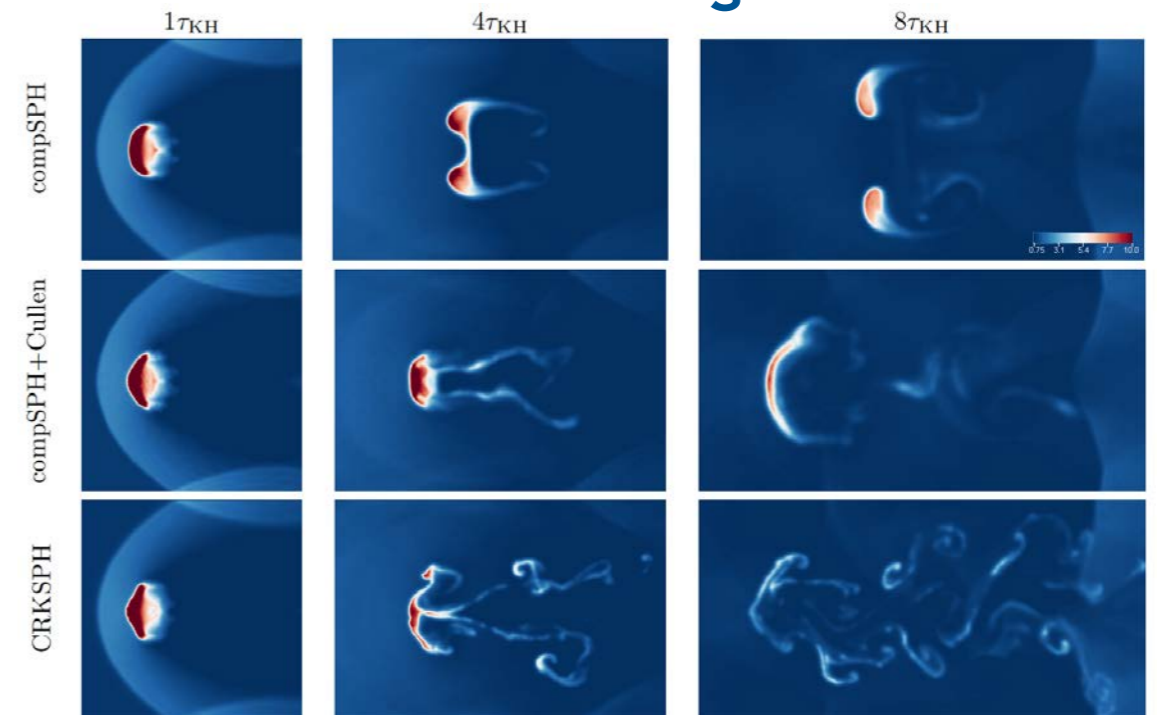


CRK-HACC: Tackling the Baryons

- HACC: Assume that the baryons follow the dark matter and the physics of the baryons is negligible on the scales of interest
- Now: Model the baryons separately as adiabatic, compressible fluid described by the Euler fluid equations; describe physics on very small scales via sub-grid models
- Range of new variables (density, pressure, temperature, ...) increases memory requirements considerably as well as computational cost
- HACC is particle-based code, therefore use Smoothed Particle Hydrodynamics (SPH) approach with new developments (Conservative reproducing kernel) to overcome traditional SPH short-comings



Triple-point shock test with CRKSPH



Blob test with CRKSPH



CRK-HACC: Performance on the KNL, PRELIMINARY!

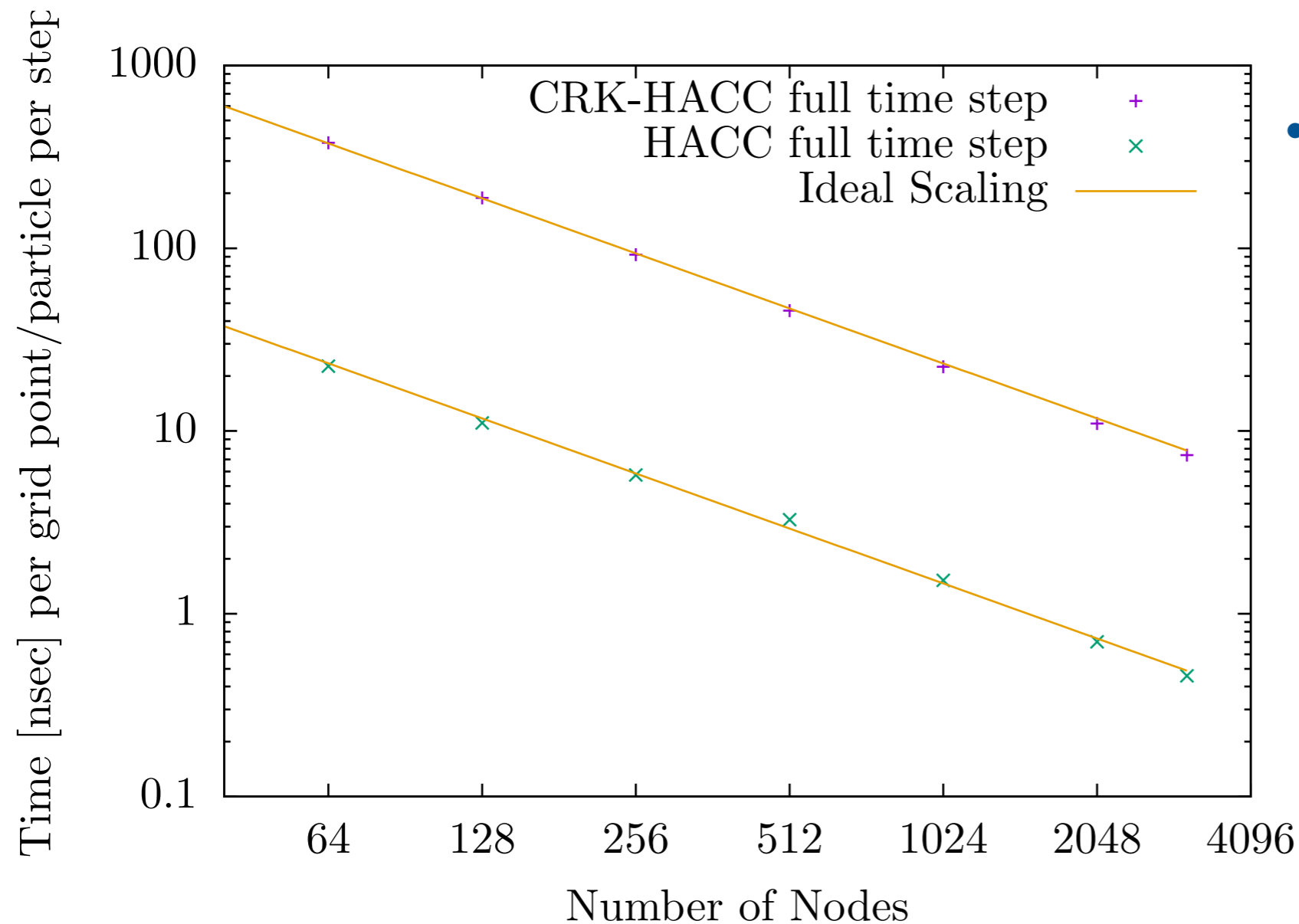
Cores	RPN	OMP	TH	Gravity [Time/s]	CRK-HACC [Time/s]	Ratio
16	4	4	16	616.33	21881	35.5
16	4	8	32	543.73	18598	34.2
16	4	16	64	530.23	19885	37.1
16	8	2	16	544.75	21750	39.9
16	8	4	32	459.53	18548	40.3
16	8	8	64	437.21	20026	45.8
16	8	4	64	468.50	21339	45.4

RPN: MPI ranks for each run; **OMP:** OpenMP threads per MPI rank, **TH:** total number of threads = RPN*OMP

- Problem set up: 2×320^3 particles, 320^3 grid, 3 full time steps, 5 sub-cycles
- Results for flat mode and cache mode are essentially identical



CRK-HACC: Strong Scaling on the KNL, PRELIMINARY!



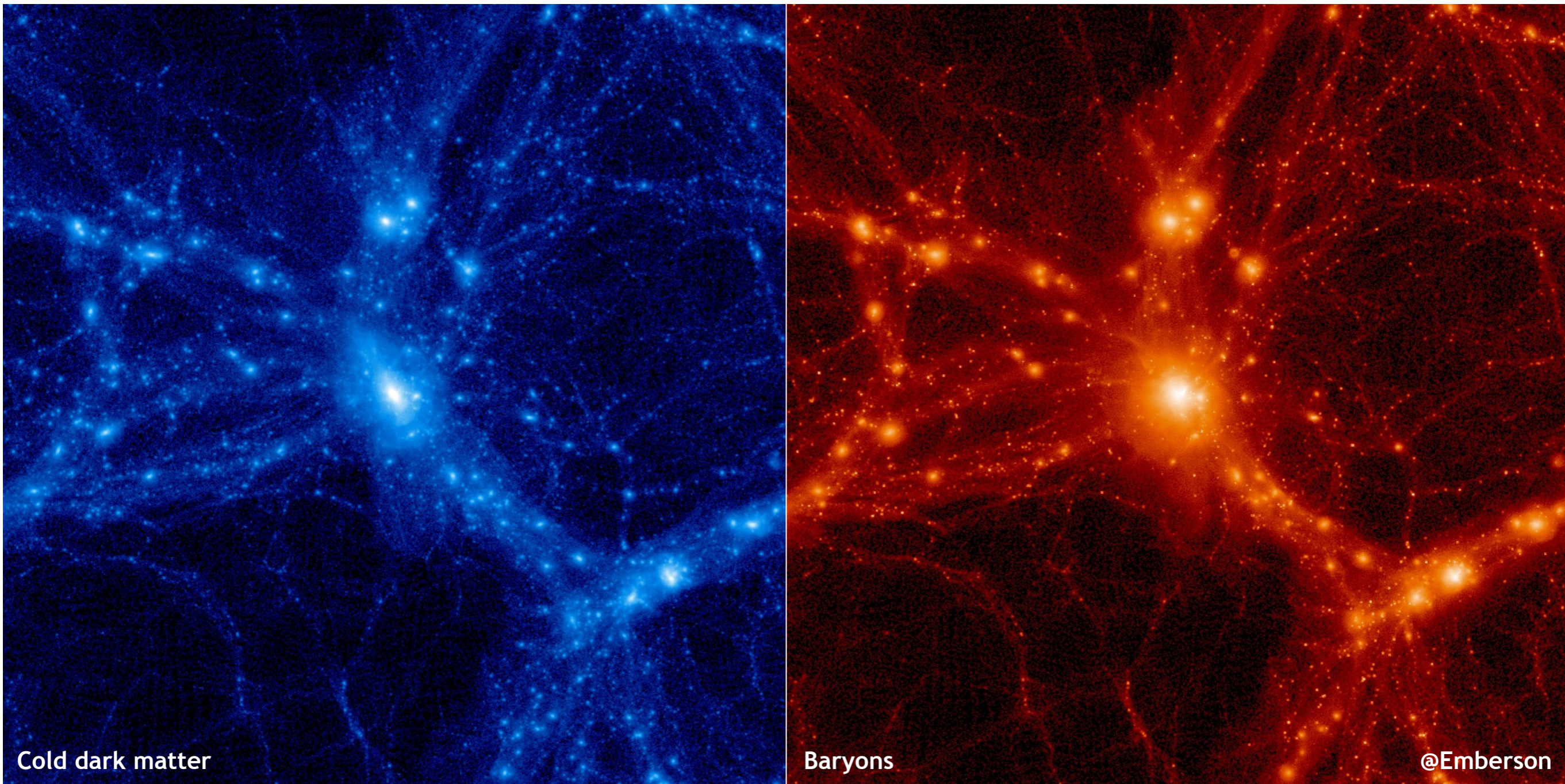
- **Problem set up:**

- ▶ 2048³ particles on 64 to 2048 nodes (and 2304³ particles on 3072 nodes)
- ▶ 8 ranks per nodes, 16 threads
- ▶ Cache-quad mode
- ▶ Per particle CRK-HACC ~15 times slower



CRK-HACC: The Santa Barbara Cluster

- Standard test for cosmological hydro codes, first carried out in 1999 (Frenck et al.)



Density slices through the full simulation volume

Final Thoughts

- HACC was designed from the start to allow easy portability between a diverse set of architectures
- Large portions of the code base (~95%) of HACC remain unchanged when moving from one machine to another; 5% are optimized in several ways (algorithms, implementation, parameters) to achieve high performance on all architectures
- HACC had been optimized for Mira already, port to Theta was straightforward
- HACC weak and strong scales on the full machine, I/O could be better
- First science run successfully carried out, currently being analyzed
- With new generation of supercomputers around the corner, we started adding new baryonic physics to HACC, now have CRK-HACC
- CRK-HACC scales also well, still needs to be optimized (implemented adaptive time stepping) and some small issues ironed out
- One specific “user” comment: the submission system could be improved (many different variables, all ALCF machines have different syntax now, and it’s different from OLCF and NERSC ... it’s really difficult to keep track of all of them!)



Computing the Sky at Extreme Scales

Why SPH?

- Galilean Invariant, intrinsically adaptive, agnostic to surface geometries, conservative, easily parallelizable/scalable, particle/interpolant connectivity is mutable, enabling more accurate modeling of extreme material deformations

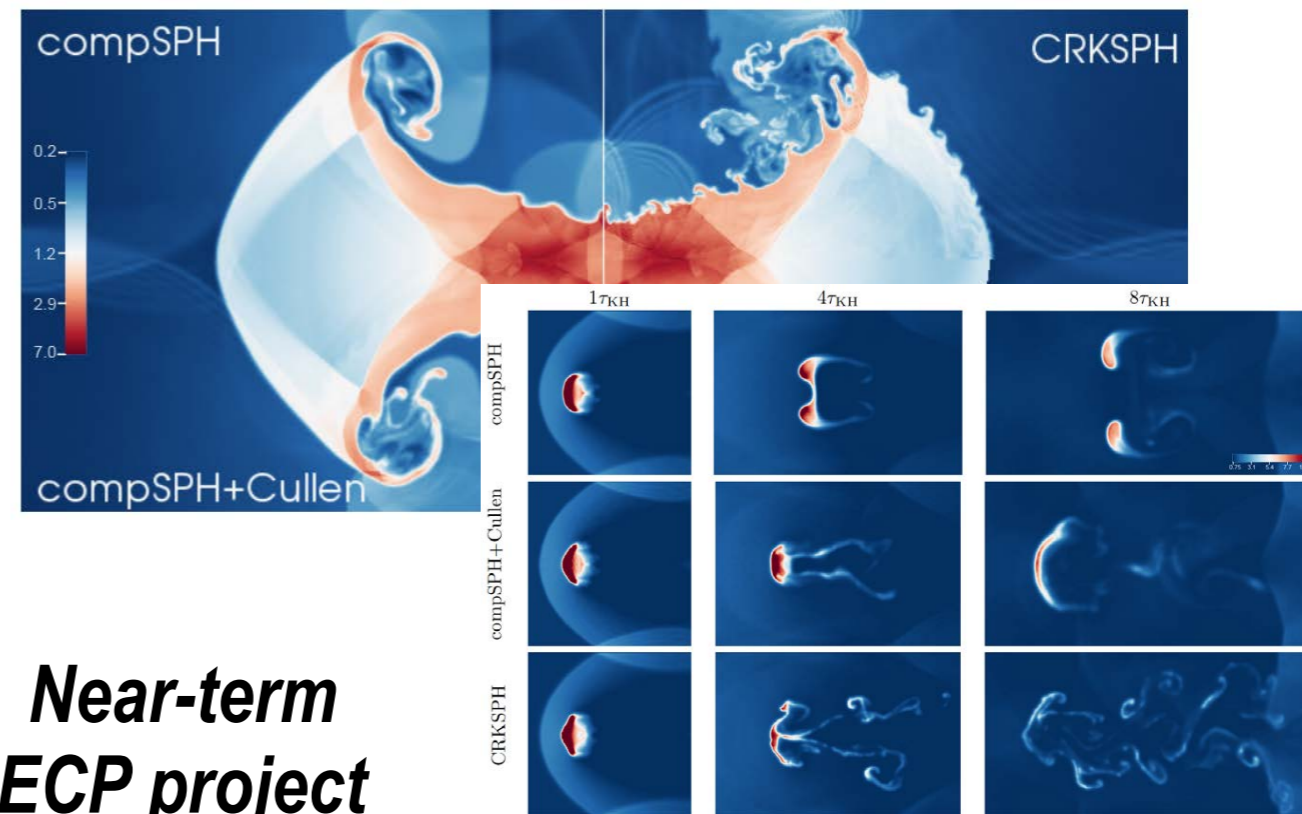
Why not?

- E0 Error: cannot reproduce even a constant field, becomes increasingly troubling over density discontinuities.
- Overly diffusive Artificial Viscosity (AV)

CRKSPH (Conservative Reproducing Kernel SPH)

- Utilizes Linear Reproducing Kernels (improves E0)
- Novel Limited Viscosity (improves AV diffusivity)
- Conservative Formalism (maintains conservation even though the kernels are no longer symmetric)

Triple-point shock test with CRKSPH



Near-term ECP project

Blob test with

*Frontiere, Raskin, Owen, J. Comp. Phys. (2017);
arXiv:1605.00725 [physics.comp-ph]*

